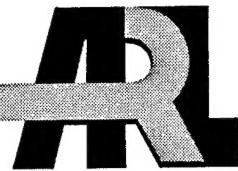


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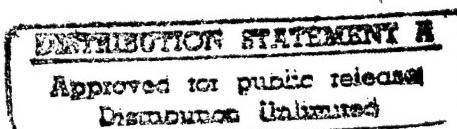


# The Asymmetry Parameter and Aggregate Particles

by  
Gorden Videen, Ronald G. Pinnick, Dat Ngo,  
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ARL-TR-1393

October 1997



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**ARL-TR-1393**

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# **The Asymmetry Parameter and Aggregate Particles**

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## Abstract

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We derive and examine the general expression for the scattering asymmetry parameter  $g$ . For aggregate particles, the asymmetry parameter is made up of two terms. One term accounts for interference effects of the electromagnetic fields radiating from the individual subsystems. The other term accounts for interaction effects of the electromagnetic fields between these subsystems. Enhanced backscatter is one phenomenon resulting from these interactions. Numerical results demonstrate that interference effects play a dominant role when the separation distance between aggregates is smaller than half the incident wavelength. As the separation distance becomes large, both interference and interaction effects drop off, and the asymmetry parameter approaches that of the individual particle constituents.

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## 1. Introduction

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The asymmetry parameter  $g$  has a long history dating back to the beginning of this century, when it was used to calculate radiation pressure exerted on particles [1]. It is currently an essential input in many radiative transfer and climate models [2]. Until recently, these models have used asymmetry parameters calculated from spherical particles, simulating water droplets and aerosol particles. Unfortunately, atmospheric particles, including ice crystals and water and aerosol particles containing contaminants, are not generally symmetric spheres. The scattering properties, including the asymmetry parameter, of these particles can vary significantly from those of spheres; e.g., both theoretical and observational work has shown that the values of the asymmetry parameter for ice crystals in cirrus clouds are significantly smaller than those given by Mie theory [3–5]. Since cirrus clouds cover about 20 to 30 percent of the earth, they influence the climate through their effects on the radiation budget [6]. A great deal of uncertainty still exists in the specification of the asymmetry parameter for cirrus clouds because of the extremely complicated shapes of ice crystals. Francis *et al.* [7] found that the deduced values of  $g$  from field experiments varied between 0.7 and 0.85. For other important radiative transfer parameters (like the extinction, scattering, and absorption efficiencies and the single scattering albedo), the anomalous diffraction approximation can provide some insight into their physical processes [8–12]. However as yet, no simple approximation can be used to calculate the asymmetry parameter.

In this report we take a closer examination of the asymmetry parameter and provide some physical insight, with the aim of stimulating further interest. We first derive the general expression for the asymmetry parameter for an arbitrary particle. Since theories have recently been developed to calculate the scatter from ensembles of particles, we consider this special case in more detail, examining specifically terms that contribute to interference and enhanced backscatter. Such theories can be used to model smoke, aerosols, and droplets containing contaminants.

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## 2. Relevant Equations

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We start by writing the general expression of the asymmetry parameter in three dimensions. We consider an arbitrarily shaped particle illuminated by a plane wave traveling in the positive  $z$  direction. For many scattering systems, such as bispheres [13–28] and spheres containing an inclusion [29–36], derivations of the scattered fields take advantage of system symmetries and the incident plane wave travels in an arbitrary direction. The appendix provides relations by which the scattering coefficients for a system illuminated by an arbitrarily incident plane wave may be converted to those for an equivalent system in which the plane wave is traveling in the positive  $z$  direction.

The scattered electric field can be expressed in terms of a vector spherical harmonic expansion as

$$\mathbf{E}^{sca} = \sum_{n=0}^{\infty} \sum_{m=-n}^n a_{nm}^{(j)} \mathbf{M}_{nm} + b_{nm}^{(j)} \mathbf{N}_{nm}, \quad (1)$$

where the index  $j$  on the scattering field coefficients  $a_{nm}$  and  $b_{nm}$  corresponds to what is acquired with incident plane-wave illumination polarized in the  $\hat{x}$  ( $j = 1$ ) and  $\hat{y}$  ( $j = 2$ ) directions. The vector spherical harmonics are defined by

$$\begin{aligned} \mathbf{M}_{nm} &= \hat{\theta} \left[ \frac{im}{\sin \theta} h_n^{(1)}(kr) \tilde{P}_n^m(\cos \theta) e^{im\phi} \right] \\ &\quad - \hat{\phi} \left[ h_n^{(1)}(kr) \frac{d}{d\theta} \tilde{P}_n^m(\cos \theta) e^{im\phi} \right], \end{aligned} \quad (2)$$

$$\begin{aligned} \mathbf{N}_{nm} &= \hat{r} \left[ \frac{1}{kr} h_n^{(1)}(kr) n(n+1) \tilde{P}_n^m(\cos \theta) e^{im\phi} \right] \\ &\quad + \hat{\theta} \left[ \frac{1}{kr} \frac{d}{dr} (rh_n^{(1)}(kr)) \frac{d}{d\theta} \tilde{P}_n^m(\cos \theta) e^{im\phi} \right] \\ &\quad + \hat{\phi} \left[ \frac{1}{kr} \frac{d}{dr} (rh_n^{(1)}(kr)) \frac{im}{\sin \theta} \tilde{P}_n^m(\cos \theta) e^{im\phi} \right], \end{aligned} \quad (3)$$

and the wavelength of the incident plane wave is  $\lambda = 2\pi/k$ ,  $h_n^{(1)}(kr)$  are the spherical Hankel functions of the first kind, and a time dependence of  $\exp(-i\omega t)$  is implicit. The normalized associated Legendre polynomials are given by

$$\tilde{P}_n^m(\cos \theta) = \sqrt{\frac{(2n+1)(n-m)!}{2(n+m)!}} P_n^m(\cos \theta). \quad (4)$$

The scattering amplitudes in the far field can be expressed by the matrix

$$\begin{pmatrix} E_{\theta}^{sca} \\ E_{\phi}^{sca} \end{pmatrix} = \frac{e^{ikr}}{-ikr} \begin{pmatrix} S_1 & S_4 \\ S_3 & S_2 \end{pmatrix} \begin{pmatrix} E_{TE}^{inc} \\ E_{TM}^{inc} \end{pmatrix}, \quad (5)$$

where  $E_{TM}^{inc}$  and  $E_{TE}^{inc}$  correspond to transverse electric and transverse magnetic incident plane wave illumination, respectively; and  $E_{\theta}^{sca}$  and  $E_{\phi}^{sca}$  correspond to scattered electric fields polarized in the  $\hat{\theta}$  and  $\hat{\phi}$  directions. In general, the elements of the scattering amplitude matrix can be written as

$$S_1 = \sum_{n=0}^{\infty} \sum_{m=-n}^n (-i)^n e^{im\phi} \times [b_{nm}^{(2)} \tilde{\pi}_n^m + a_{nm}^{(2)} \tilde{\tau}_n^m], \quad (6)$$

$$S_2 = -i \sum_{n=0}^{\infty} \sum_{m=-n}^n (-i)^n e^{im\phi} \times [a_{nm}^{(1)} \tilde{\pi}_n^m + b_{nm}^{(1)} \tilde{\tau}_n^m], \quad (7)$$

$$S_3 = -i \sum_{n=0}^{\infty} \sum_{m=-n}^n (-i)^n e^{im\phi} \times [a_{nm}^{(2)} \tilde{\pi}_n^m + b_{nm}^{(2)} \tilde{\tau}_n^m], \text{ and} \quad (8)$$

$$S_4 = \sum_{n=0}^{\infty} \sum_{m=-n}^n (-i)^n e^{im\phi} \times [b_{nm}^{(1)} \tilde{\pi}_n^m + a_{nm}^{(1)} \tilde{\tau}_n^m], \quad (9)$$

where

$$\tilde{\pi}_n^m = \frac{m}{\sin \theta} \tilde{P}_n^m(\cos \theta), \text{ and} \quad (10)$$

$$\tilde{\tau}_n^m = \frac{\partial}{\partial \theta} \tilde{P}_n^m(\cos \theta). \quad (11)$$

The asymmetry parameter is the integral of the cosine-weighted intensity phase function and is a measure of the amount of forward scatter from the particle:

$$g = \frac{1}{2k^2 C_{sca}} \int_0^{2\pi} \int_0^\pi (|S_1|^2 + |S_2|^2 + |S_3|^2 + |S_4|^2) \cos \theta \sin \theta d\theta d\phi, \quad (12)$$

where the scattering cross section is defined as

$$C_{sca} = \frac{1}{2k^2} \int_0^{2\pi} \int_0^\pi (|S_1|^2 + |S_2|^2 + |S_3|^2 + |S_4|^2) \sin \theta d\theta d\phi, \quad (13)$$

which is equivalent to

$$C_{sca} = \frac{2\pi}{k^2} \sum_{n=1}^{\infty} n(n+1) \sum_{m=-n}^n \left( |a_{nm}^{(1)}|^2 + |b_{nm}^{(1)}|^2 + |a_{nm}^{(2)}|^2 + |b_{nm}^{(2)}|^2 \right). \quad (14)$$

For an arbitrarily shaped particle, there may be coupling between the modes, and the scattering amplitude matrix elements  $S_3$  and  $S_4$  are not zero. After a fair amount of algebraic manipulations, the asymmetry parameter may be expressed as

$$\begin{aligned} g = & \frac{4\pi}{k^2 C_{sca}} \operatorname{Re} \sum_{n,m} m \left( a_{nm}^{(1)} b_{nm}^{(1)*} + a_{nm}^{(2)} b_{nm}^{(2)*} \right) \\ & + i n (n+2) \sqrt{\frac{(n-m+1)(n+m+1)}{(2n+1)(2n+3)}} \\ & \times \left( a_{nm}^{(1)} a_{n+1m}^{(1)*} + b_{nm}^{(1)} b_{n+1m}^{(1)*} + a_{nm}^{(2)} a_{n+1m}^{(2)*} + b_{nm}^{(2)} b_{n+1m}^{(2)*} \right). \quad (15) \end{aligned}$$

Equation (15) contains a great deal of information. If, for instance, there is no correlation between the coefficients,

$$\langle a_{nm}^{(j)} a_{n+1m}^{(j)*} \rangle = \langle b_{nm}^{(j)} b_{n+1m}^{(j)*} \rangle = \langle a_{nm}^{(j)} b_{nm}^{(j)*} \rangle = 0, \quad (16)$$

then there is no preferential scattering hemisphere. This condition is necessary for a scatterer to be Lambertian.

Equation (15) also shows that there is no preferential scattering hemisphere for any individual mode  $\mathbf{M}_{nm}$  or  $\mathbf{N}_{nm}$ . However, when more than one mode is considered, there can be preferential scattering. A preferential scattering hemisphere is the result of interference between the modes. In general, constructive interference tends to be in the forward-scatter region, leading to positive  $g$ . This phase information (and resulting scattering directionality) is contained within the incident plane wave coefficients.

For reference, we show in figure 1 the calculated asymmetry parameter as a function of radius for carbon and water spheres. For radii that are small compared to the wavelength, the asymmetry parameter is zero, since no preferential scattering hemisphere exists for isolated Rayleigh-size scatterers. As the sphere radius approaches the wavelength, the asymmetry parameter rises rapidly, corresponding to the increase in scatter in the forward direction. For the highly absorbing carbon spheres, the asymmetry parameter levels off around 0.9. As the transparent, water-sphere radius increases, the internal fields selectively intensify specific modes in the sphere through morphology-dependent resonances. Selective enhancement of specific modes tends to decrease the asymmetry parameter, since (from eq (16)) a specific mode does not have a preferential scattering hemisphere, and the normalizing scattering cross section is increased.

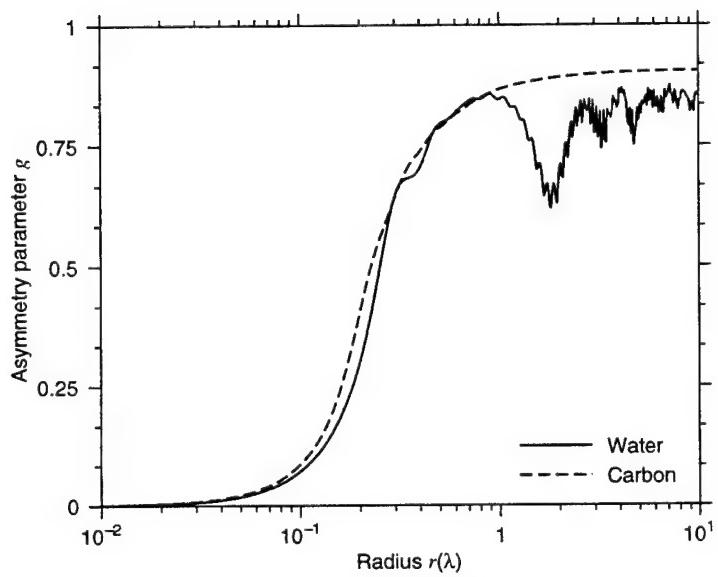


Figure 1. Asymmetry parameter of water ( $n = 1.33$ ) and carbon ( $n = 1.75 + 0.44i$ ) spheres as a function of sphere radius.

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### 3. Aggregates

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Other researchers [13–28] have recently derived solutions for the scatter from aggregate particles by considering the entire system to be composed of multiple subsystems, and by including an interaction term as part of the field incident on each subsystem. This interaction term is due to the scattered field from all other subsystems striking the subsystem of interest. The scattering coefficients for the  $j$ th subsystem,  $f_{nm}^{1,j}$  and  $f_{nm}^{2,j}$ , can be expressed as a set of coupled linear equations,

$$f_{nm}^{1,j} = a_{nm}^{1,j} + \sum_{k \neq j} \sum_{n'} \sum_{m'} f_{n'm'}^{1,k} A_{nm}^{jkn'm'} + f_{n'm'}^{2,k} B_{nm}^{jkn'm'}, \quad (17)$$

$$f_{nm}^{2,j} = a_{nm}^{2,j} + \sum_{k \neq j} \sum_{n'} \sum_{m'} f_{n'm'}^{1,k} C_{nm}^{jkn'm'} + f_{n'm'}^{2,k} D_{nm}^{jkn'm'}, \quad (18)$$

where  $a_{nm}^{n,j}$  is a function of the incident field and  $A_{nm}^{jkn'm'}$ ,  $B_{nm}^{jkn'm'}$ ,  $C_{nm}^{jkn'm'}$ , and  $D_{nm}^{jkn'm'}$  are system-dependent parameters that include translation coefficients that can be used to express the vector spherical harmonics in subsystem  $k$  in terms of vector spherical harmonics in subsystem  $j$ . Once solutions for the scattering coefficients  $f_{nm}^{1,j}$  and  $f_{nm}^{2,j}$  are found, the total scattered field is found as the superposition of the scattered fields from all the subsystems. We can find the asymmetry parameter using equation (15) by expressing the total scattered field in terms of a single coordinate system, which is a straightforward process using the addition theorem for vector spherical harmonics and the linearity of the system.

Aggregation can affect the scatter in two ways: through interference and interaction. We can isolate the effect of interference by expressing the total scattered fields in terms of the individual scattering components. Since the total scattered electric field can be expressed as the sum of the contributions from each aggregate ( $\mathbf{E}^{sca} = \sum_j \mathbf{E}_j^{sca}$ ), the asymmetry parameter can be expressed as

$$\begin{aligned} g &= \frac{\sum_{j \neq k} \int_0^{2\pi} \int_0^\pi \mathbf{E}_j^{sca*} \cdot \mathbf{E}_k^{sca} \cos \theta \sin \theta d\theta d\phi}{\int_0^{2\pi} \int_0^\pi \mathbf{E}^{sca*} \cdot \mathbf{E}^{sca} \sin \theta d\theta d\phi} \\ &\quad + \frac{\sum_j \int_0^{2\pi} \int_0^\pi \mathbf{E}_j^{sca*} \cdot \mathbf{E}_j^{sca} \cos \theta \sin \theta d\theta d\phi}{\int_0^{2\pi} \int_0^\pi \mathbf{E}^{sca*} \cdot \mathbf{E}^{sca} \sin \theta d\theta d\phi}. \end{aligned} \quad (19)$$

The first term of equation (19) accounts for the interference between the fields scattered by the  $j$ th and  $k$ th subsystems. The second term of equation (19) is the sum of the individual asymmetry parameters calculated from each individual subsystem. In sections 3.1 and 3.2, we consider these terms individually. When the subsystems are far enough apart that any particle-particle interaction or interference becomes negligible, equation (19) must reduce to

$$g_o = \frac{\sum_j g_j C_{sca,j}}{C_{sca}}, \quad (20)$$

where  $g_j$  and  $C_{sca,j}$  are the asymmetry parameter and scattering cross section for the isolated  $j$ th subsystem. The effects of the interaction are contained within the scattering coefficients themselves; thus, we can isolate the interaction effect by subtracting  $g_o$  from the second term of equation (19).

### 3.1 Interaction

Interaction between subsystems is difficult to characterize because there is mode mixing between the subsystems. One effect of the interaction is enhanced backscatter, which is due to constructive interference of rays reflecting off multiple interfaces. In the backscatter direction, the path difference of light rays striking multiple interfaces is the same when the order of the interfaces is reversed. These two rays (forward and backward traversing) interfere constructively, and the resulting intensity is enhanced. When backscattered light is enhanced, the asymmetry parameter must decrease.

Figure 2 shows the individual components of the asymmetry parameter given by equation (19) as a function of separation distance  $d$ . In this figure, the asymmetry parameter is averaged over all orientations of a scattering system composed of two  $r = \lambda/2$  carbon spheres (fig. 2a) and two  $r = \lambda/10$  carbon spheres (fig. 2b). The effect of enhanced backscatter is contained within the second term of equation (19) (after  $g_o$  is subtracted out). Note that the interference term is always positive, and the interaction term is always negative. For both particle systems, the interaction component is approximately proportional to  $d^{-2}$ , which is proportional to the scattered flux of one subsystem, intercepted by the other subsystem. This is to be expected, since the interaction between particles must decrease with the magnitude of the interaction field.

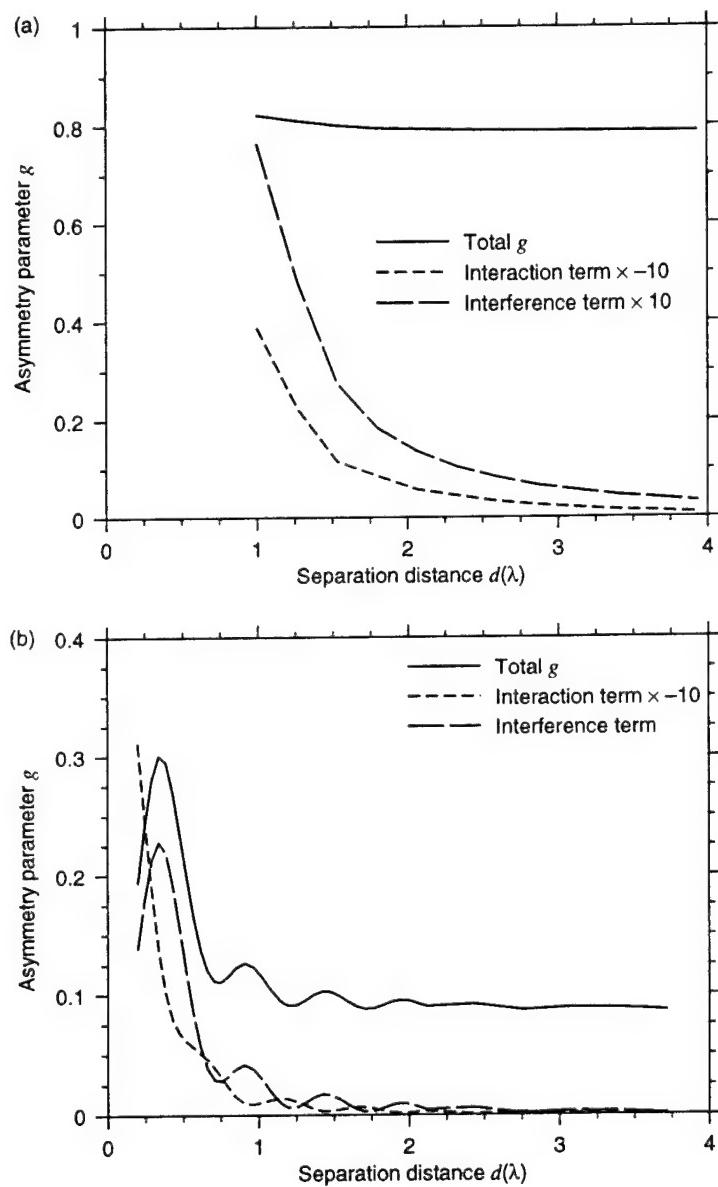


Figure 2. Asymmetry parameter as a function of separation distance  $d$  for two carbon spheres, (a)  $r = \lambda/2$  and (b)  $r = \lambda/10$ , averaged over all orientations. Note interference term is always positive and interaction term is always negative.

### 3.2 Interference

Interference produces a high-frequency structure on the scattering phase function. This is illustrated in figure 3, which shows the scatter from a pair of  $r = \lambda/4$  carbon spheres illuminated at broadside incidence. For comparison, the scatter from a single sphere having the same parameters is also shown. The major effect of interference is a high-frequency modulation on the scattered signals. The positions of the high-frequency maxima and minima are dependent on the positions of the aggregate subsystems. Since the scattered waves from each subsystem acquire the same phase difference in the forward direction, there is always constructive interference in this direction, and therefore, a maximum. Hence, interference tends to increase the asymmetry parameter. This is illustrated in figure 2a, which shows the individual components of the asymmetry parameter (eq (19)) as a function of separation distance  $d$ . Note that as the separation distance  $d$  increases, the interference component of the asymmetry parameter for the two-sphere aggregate is positive and asymptotically approaches 0. This is because the spatial frequency of the interference structure increases with separation distance, and a high-frequency modulation has little effect on the integration of the scattering intensities given by equation (19). In figure 2b (for  $r = \lambda/10$  carbon spheres), there is much more structure in the interference component. The large maximum occurring near  $d = \lambda/4$  is due to destructive interference of scattered light in the backward direction, which increases the

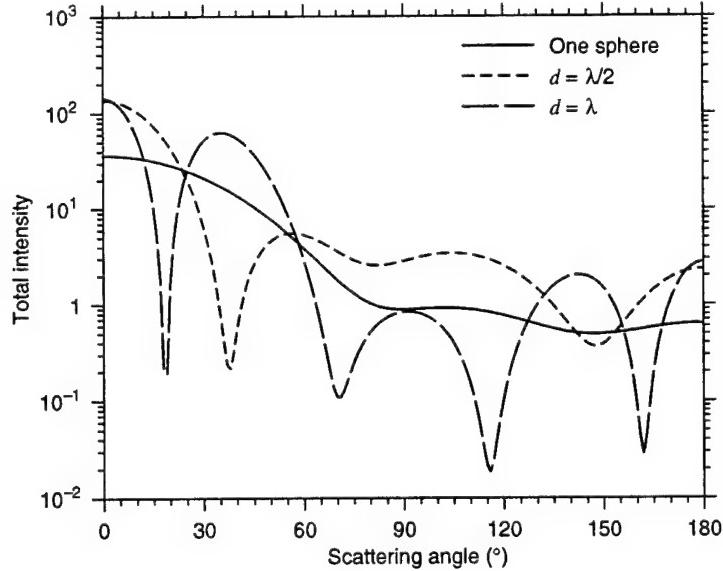


Figure 3. Intensity as a function of scattering angle for  $r = \lambda/4$  carbon spheres ( $n = 1.75 + 0.44i$ ) showing interference structure.

value of  $g$ . Subsequent maxima in the interference term occur at intervals of approximately  $\lambda/2$  and are similar to the interference structure of thin films. The particulate components of a simple aggregate tend to be in contact. Figure 4a shows the asymmetry parameter as a function of radius  $r$  for two carbon spheres in contact, averaged over all orientations. The curves of this figure are significantly different from the curves of figure 2, owing to the spheres being in contact. All that remains of the interference structure is the large maximum occurring near  $r = \lambda/8$  ( $d = \lambda/4$ ), which dominates the asymmetry parameter for small sphere radii. The interference term of equation (19) remains a significant component for much larger separation distances when the spheres are in contact. The shape of the total asymmetry parameter is similar to that of the single carbon sphere (shown for comparison in fig. 1).

Figure 4b shows the percentage errors resulting in asymmetry parameter calculations when the two-sphere aggregate is assumed to be a sphere. Although equivalent-volume spheres approximate the asymmetry parameter better than equivalent-area spheres and an isolated sphere having the same radius of one of the components, the errors are still significant, especially when the aggregate components are smaller than the wavelength. As shown in figure 4a, at these small radii, the interference between the individual particles (which cannot be included in any equivalent-sphere system) plays a dominant role in the determination of the asymmetry parameter.

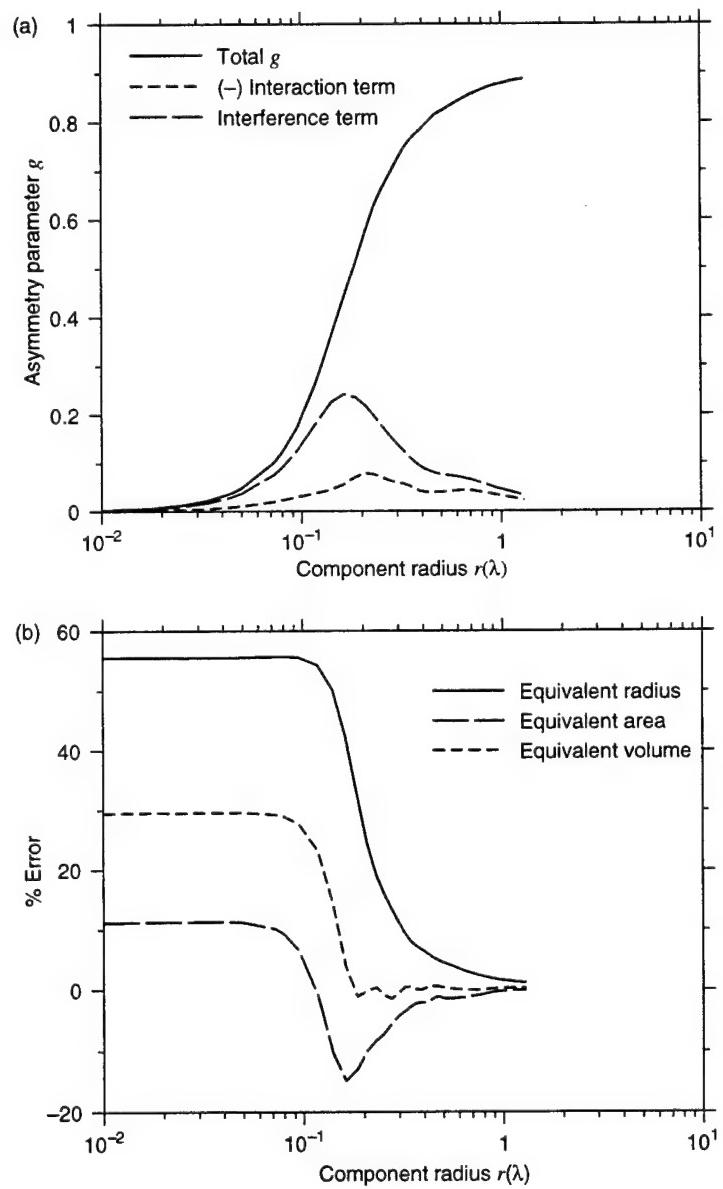


Figure 4. Asymmetry parameter: components as a function of component radius  $r$  (a) for two carbon spheres in contact, averaged over all orientations, and (b) for two-sphere aggregate compared with equivalent single carbon sphere.

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#### 4. Conclusion

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In this report, we provide an expression for the asymmetry parameter of an arbitrarily shaped particle for which the scattering solution can be expressed in a multipole solution. We have considered the special case of an aggregate and have isolated the two mechanisms—interference and interaction—that affect the value of the asymmetry parameter. Interference tends to increase the asymmetry parameter by preferentially scattering light in the forward direction. Counteracting this effect is the interaction component, which accounts for enhanced backscatter.

For the two-sphere aggregates we examined, the interference effects outweigh the interaction effects. Equivalent-sphere systems do not have a mechanism to reproduce either of these effects and, consequently, provide poor approximations when these effects are significant. Although interference plays a greater role than interaction for the bisphere system, we cannot assume that it does so for other aggregate or irregular systems. As the number of surface irregularities increases (for instance, in a multifaceted ice crystal), the interaction would be expected to play a dominant role and the asymmetry parameter would be reduced. By isolating the factors that affect the asymmetry parameter, we can proceed in the parameterization of these factors as a function of the number of subparticles contained in a system and the number of surface irregularities. Such a parameterization is immediately applicable to nonspherical ice crystals in cirrus clouds.

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## Appendix. Relations for Scattering Coefficients of Two Rotated Coordinate Systems

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The relations for the scattering coefficients of two rotated coordinate systems have been derived by Stein.<sup>1</sup> The vector spherical harmonics in the original (primed) coordinate system are related to the vector spherical harmonics in a rotated system (in which the plane wave travels in the positive  $z$  direction):

$$\mathbf{M}_{nm}^{(3)} = \sum_{m'} D_{m'}^{(n,m)} \mathbf{M}_{nm'}^{(3)'}, \quad (\text{A-1})$$

$$\mathbf{N}_{nm}^{(3)} = \sum_{m'} D_{m'}^{(n,m)} \mathbf{N}_{nm'}^{(3)'}, \quad (\text{A-2})$$

where the rotation coefficients  $D_{m'}^{(n,m)}$  are given by

$$\begin{aligned} D_{m'}^{(n,m)} &= \exp [i(m'\alpha + m\gamma)] \left[ \frac{(n+m')! (n-m')!}{(n+m)! (n-m)!} \right]^{1/2} \\ &\times \sum_{\sigma} \binom{n+m}{n-m'-\sigma} \binom{n-m}{\sigma} (-1)^{n+m-\sigma} [\cos(\beta/2)]^{2\sigma+m'+m} \\ &\times [\sin(\beta/2)]^{2n-2\sigma-m'-m} \end{aligned} \quad (\text{A-3})$$

and  $\alpha$ ,  $\beta$ , and  $\gamma$  are Euler angles following the convention of Edmonds.<sup>2</sup> The coefficients in the rotated coordinate system,  $a_{nm}$  and  $b_{nm}$ , can be expressed in the unrotated coordinate system,  $c_{nm}$  and  $d_{nm}$ , as

$$a_{nm} = \sum_{m'} D_{m'}^{(n,m)} c_{nm'}, \quad (\text{A-4})$$

$$b_{nm} = \sum_{m'} D_{m'}^{(n,m)} d_{nm'}. \quad (\text{A-5})$$

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<sup>1</sup>S. Stein, "Addition theorems for spherical wave functions," *Q. Appl. Math.* **19** (1961), 15–24.

<sup>2</sup>A. R. Edmonds, *Angular Momentum in Quantum Mechanics*, Princeton University Press, Princeton, NJ (1957).

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## REPORT DOCUMENTATION PAGE

*Form Approved  
OMB No. 0704-0188*

Public reporting burden for this collection of information is estimated to average 1 hour per response, including the time for reviewing instructions, searching existing data sources, gathering and maintaining the data needed, and completing and reviewing the collection of information. Send comments regarding this burden estimate or any other aspect of this collection of information, including suggestions for reducing this burden, to Washington Headquarters Services, Directorate for Information Operations and Reports, 1215 Jefferson Davis Highway, Suite 1204, Arlington, VA 22202-4302, and to the Office of Management and Budget, Paperwork Reduction Project (0704-0188), Washington, DC 20503.

1. AGENCY USE ONLY <i>(Leave blank)</i>		2. REPORT DATE October 1997	3. REPORT TYPE AND DATES COVERED Progress, from 1 Oct 1996 to April 1997
4. TITLE AND SUBTITLE  The Asymmetry Parameter and Aggregate Particles			5. FUNDING NUMBERS  DA PR: B53A PE: P61102
6. AUTHOR(S)  Gorden Videen, Ronald G. Pinnick (ARL), Dat Ngo (Ngo Co.), Qiang Fu, and Petr Chýlek (Dalhousie University)			
7. PERFORMING ORGANIZATION NAME(S) AND ADDRESS(ES)  U.S. Army Research Laboratory Attn: AMSRL-IS-EE 2800 Powder Mill Road Adelphi, MD 20783-1197			8. PERFORMING ORGANIZATION REPORT NUMBER  ARL-TR-1393
9. SPONSORING/MONITORING AGENCY NAME(S) AND ADDRESS(ES)  U.S. Army Research Laboratory 2800 Powder Mill Road Adelphi, MD 20783-1197			10. SPONSORING/MONITORING AGENCY REPORT NUMBER
11. SUPPLEMENTARY NOTES  AMS code: 611102.53A11 ARL PR: 7FEJ60			
12a. DISTRIBUTION/AVAILABILITY STATEMENT  Approved for public release; distribution unlimited.			12b. DISTRIBUTION CODE
13. ABSTRACT <i>(Maximum 200 words)</i>  We derive and examine the general expression for the scattering asymmetry parameter $g$ . For aggregate particles, the asymmetry parameter is made up of two terms. One term accounts for interference effects of the electromagnetic fields radiating from the individual subsystems. The other term accounts for interaction effects of the electromagnetic fields between these subsystems. Enhanced backscatter is one phenomenon resulting from these interactions. Numerical results demonstrate that interference effects play a dominant role when the separation distance between aggregates is smaller than half the incident wavelength. As the separation distance becomes large, both interference and interaction effects drop off, and the asymmetry parameter approaches that of the individual particle constituents.			
14. SUBJECT TERMS  Scatter, radiation transfer, cirrus clouds			15. NUMBER OF PAGES 28
			16. PRICE CODE
17. SECURITY CLASSIFICATION OF REPORT Unclassified	18. SECURITY CLASSIFICATION OF THIS PAGE Unclassified	19. SECURITY CLASSIFICATION OF ABSTRACT Unclassified	20. LIMITATION OF ABSTRACT UL